

PRELIMINARY QUALITY ANALYSIS OF GF-7 SATELLITE LASER ALTIMETER FULL WAVEFORM DATA

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ABSTRACT:

The full-waveform data is the core data of GaoFen-7 (GF-7) satellite laser altimeter, and evaluation of waveform data quality is an important step and premise for satellite laser altimetry and quality control. In this paper, the full waveform data quality assessment and analysis of GF-7 laser altimeter is implemented during the period of on-orbit experiment, and the real waveform data of many orbits is used to quantitatively describe the characteristic parameters of the transmitted waveform and the signal-to-noise ratio (SNR), and the result of two beam lasers is compared. The conclusion is validated that the GF-7 laser altimeter can obtain effective waveform data and the echo waveform availability of the experimental data is approximate 72.59%, moreover, the quality of beam 1 is slightly better than that of the beam 2. The laser temperature is an important indication of the quality of transmitted waveform according to the SNR changing. The good SNR value of the waveform and small footprint size will be helpful for the terrain information extraction and analysis, although the repetition frequency is low.

1. INTRODUCTION

As an active remote sensing, laser altimeter has been concerned to load on the satellite to obtain the high precision elevation information (Tang et al., 2017). The United States has launched two laser altimetry satellites, ICESat (Ice, Cloud and land Elevation Satellite) and ICESat-2, in 2003 and 2018 respectively, and based on the laser altimetry data collected by these two satellites, especially the waveform data of ICESat (Brenner et al., 2011), much research and application has been carried out, including the establishment of the theoretical model of laser altimeter, the decomposition method of echo waveform (Li et al., 2019), filtering method of the waveform (Wang et al., 2019), monitoring of the water level in lakes (Wang et al., 2012; Zhang et al., 2011), analysis of elevation changes in polar ice caps (Kääb et al., 2012; Bolch et al., 2013), estimation of forest biomass (Xing et al., 2010; Harding and David, 2005), and acquisition of ground high precision control points (Carabajal et al., 2010; Li et al., 2017).

In 2016, China launched the ZY3-02 satellite, which was equipped with the first experimental laser altimeter in China (Li et al., 2017). It was mainly used to test the function and performance of the laser altimeter, and has provided important reference for the next laser altimetry satellites although it did not record the full-waveform data. GF-7 Satellite (GF-7) is the first civilian submeter-class optical transmission stereo mapping satellite of China, which has been successfully launched on November 3rd, 2019. Along with high resolution stereo cameras, a two-beam laser altimeter with full waveform sampling and recording function is also equipped on the GF-7 to obtain high accuracy elevation control points.

In this paper, a total of 10 orbits laser altimetry data obtained by GF-7 satellite during the period of the on-orbit test are used to evaluate the waveform data quality. The waveform characteristic

parameters, SNR (signal-to-noise ratio) and other relative parameters are calculated and compared in the experiment, and the results are analysed to summarize the evaluation results of the waveform data quality of GF-7 laser altimeter.

2. GF-7 SATELLITE AND LASER ALTIMETER

2.1 The basic parameters of GF-7

The GF-7 satellite has two stereo camera and two laser beam, illustrated in Figure 1. The viewing angle of forward and backward camera is 26° and -5° respectively, and the swath width is about 20km with resolution better than 0.8m.

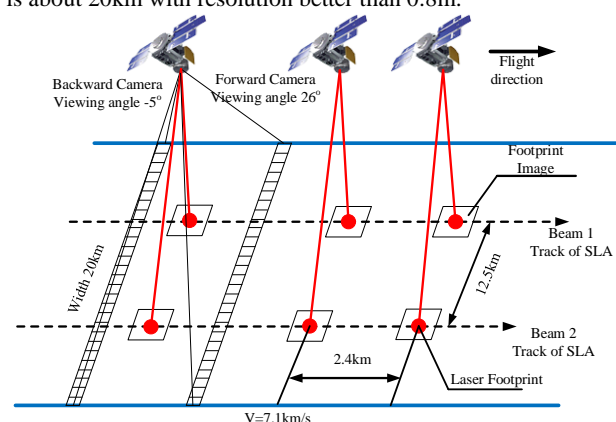


Figure 1. Working mode of GF-7 satellite

And the inclination angle of two laser beams are about 0.7 degrees on the vertical plane along the left and right of nadir pointing direction respectively, and the laser wavelength is 1064nm, with a divergence angle of [30urad,35urad], which is better than the

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designed value 60urad and corresponds to surface laser spot in a diameter of [15m,17.5m]@500km. The sampling interval of the full-waveform is 0.5ns, and the corresponding distance resolution is 0.075m. In order to adapt to the echo intensity of difference surface reflectance, the full-waveform data are recorded in two kind of modes-high and low gain, which are showed in Table 1.

Parameter	Value
Wavelength	1064nm
Beam number	2
Frequency	3Hz/6Hz
Laser divergence angle	[30urad,35urad]
Footprint diameter	[15m-17.5m]@500km
Waveform Sampling interval	2GHz/0.5ns
Waveform quantization bits	10bits
Waveform sampling mode	High/Low gain
Pulse width of transmitting waveform	5-7ns
Footprint image resolution	3.2m
Footprint image size	550*550pixel

Table. 1 Parameters of GF-7 satellite laser altimeter

2.2 Experimental data

In this paper, a total number of 10-orbit laser data obtained during the period of on-orbit test of GF-7 were statistically analysed. These 10-orbit laser waveform data were collected, and the laser footprints are distributed in some regions of North America, Europe, Africa, Oceania, Asia and a few of Antarctica. The number of laser points in each orbit is shown in Table 2, and when the proportion of effective echo is low, it means that laser propagation is blocked by the thick cloud, with no echo waveform detected by the receiver.

Orbit No.	Number of Transmitted laser pulses	Number of Received laser pulses	Ratio of effective echoes	Acquisition time
117	1224	469	38.32%	2019.11.11
154	1202	958	79.70%	2019.11.13
245	750	746	99.47%	2019.11.19
246	682	615	90.18%	2019.11.19
282	1890	1669	88.31%	2019.11.22
320	1350	1317	97.56%	2019.11.24
321	1442	541	37.52%	2019.11.24
337	1606	1553	96.70%	2019.11.25
344	948	223	23.52%	2019.11.26
358	1652	1161	70.28%	2019.11.26
Total	12746	9252	72.59%	

Table 2. Orbital number, number of laser points and acquisition time of experimental data

3. WAVEFORM QUALITY EVALUATION METHOD

3.1 Waveform background noise parameters

The noise of waveform data is mainly composed of the system noise of the laser itself and the solar background noise. And different terrain features may impose different influence on the waveform, here, we make an assessment merely on the system noise of the laser.

The background noise was calculated from the intensity of the sampling point in the area with no effective waveform in the GF-

7 waveform data. The transmitted waveform length of GF-7 laser altimeter is 200ns, and a total number of 400 sampling points are obtained with a sampling frequency of 2GHz. The last quarter of transmitted waveform, that is, beginning from 150ns, sampling points of 50ns are used to calculate the waveform background noise. The received waveform length is 400ns, corresponding to 800 sampling points, and the last 50ns waveform data are used to calculate the background noise of received waveform

$$m_N = \sum_t^T \frac{V_i}{(T-t)} \quad (1)$$

$$\sigma_N = \sqrt{\sum_t^T \frac{(V_i - m_N)^2}{(T-t)^2 - 1}} \quad (2)$$

Where t and T are beginning and ending time coordinates of the range respectively. For the transmitted waveform, the $t = 150ns$ and $T = 200ns$. The V_i is the intensity value of the waveform at the moment t . And for the receive waveform, $t = 350ns$ and $T = 400ns$.

3.2 Waveform gaussian characteristic parameters

The waveform data of the satellite laser altimeter is approximately a gaussian function, which is used to fit the waveform to obtain the gaussian characteristic parameters, such as amplitude value of waveform. The Gaussian model formula is as follows,

$$V(t) = \sum_{i=1}^N A_i e^{-\frac{(t-\mu_i)^2}{2\sigma_i^2}} \quad (3)$$

Where A_i is the amplitude value of i -th gaussian component, μ_i is the peak location of the i -th gaussian component, and σ_i is the root mean square (RMS) pulse width of the i -th gaussian component.

Firstly, the effective waveform is selected out using the above background noise value plus four times the standard deviation as the threshold, and the gaussian fitting is carried out, using the nonlinear least square method for iteration to ensure the fitting accuracy. Finally, the gaussian parameters are extracted.

3.3 Waveform kurtosis and skewness

For the transmitted waveform, the shape should be almost the same and viewed as one Gaussian function distribution during the working period. The Kurtosis of the waveform is an indicator to reflect the sharp or flat degree at the top of distribution curve of waveform intensity, and the skewness of the waveform is a measure of the skew direction and degree of the statistical data distribution, which can directly show the asymmetric degree of the data distribution. The arithmetic mean, standard deviation values and skewness coefficients of two different data sets sometimes may be same, however, the sharp degree may be different. The Kurtosis and Skewness of a waveform data set with n samples is as follows,

$$SK = \frac{n \sum_{i=1}^n (x_i - \bar{x})^3}{(n-1)(n-2)SD^3} \quad (4)$$

$$K = \frac{1}{nSD^2} \sum_{i=1}^n (x_i - \bar{x})^4 \quad (5)$$

Where n is the total number of samples, and SD is the variance of the waveform as $SD = \frac{1}{n} \sum_{i=1}^n (x_i - \bar{x})^2$ (\bar{x} is the mean value of waveform samples).

The Kurtosis value of a standard gaussian distribution is 3 ($K = 3$). When $K > 3$, there is a sharp peak on top of the statistical data distribution curve, indicating that the data is more concentrated. And conversely, there is a flat peak of the data distribution curve corresponding to a more dispersed distribution. The Skewness value of the standard gaussian is 0 ($SK = 0$). When $SK > 0$, it can be considered that the statistical data has a right-skewed distribution, that is, it shows a trailing distribution on the right, and the greater the SK is, the higher degree of right

deviation is. Conversely, the statistical data has a left-skewed distribution with $SK < 0$, showing a trailing distribution on the left, and the smaller the SK is, the higher the degree of left deviation is.

3.4 Waveform signal-to-noise ratio (SNR)

The SNR of waveform is calculated based on the mean value and standard deviation of background noise (Nie et al.,2014; Li, 2017), and the formula is as follows,

$$SNR = 10lg \frac{I_{max} - m_N}{\sigma_N} \quad (6)$$

Where the I_{max} is maximum value of the waveform intensity, and m_N is the mean value of background noise, σ_N is the

standard deviation of the waveform noise. This method does not need to fit the waveform to get a more accurate SNR, and is computational convenient.

4. RESULTS AND ANALYSIS

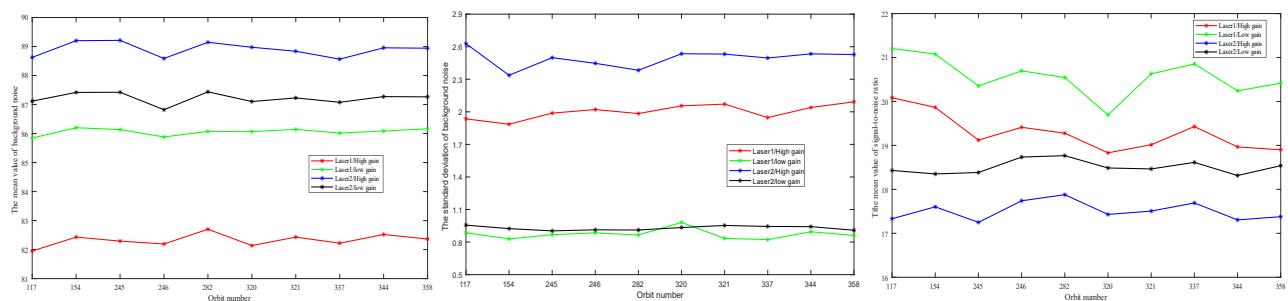
4.1 Results of the transmitted waveform experiment

According to the above-mentioned quality evaluation parameters, the results of the experiments are shown in Table 3 and Figure 2.

Orbit Number	Laser	Gain mode	Background noise (mV)		SNR	Amplitude (mV)	Peak Location(ns)	Pulse width(ns)	Skewness	Kurtosis
			Mean	Std						
117	Laser1	HG*	81.951	1.935	20.084	112.220	217.053	6.772	0.168	2.705
		LG*	85.846	0.886	21.204	65.915	207.093	6.240	0.482	2.644
	Laser2	HG	88.625	2.628	17.332	77.895	215.118	6.054	0.060	2.678
		LG	87.117	0.956	18.431	36.173	205.197	5.207	0.411	2.592
154	Laser1	HG	82.430	1.886	19.866	103.932	216.837	6.867	0.144	2.701
		LG	86.200	0.829	21.078	60.568	206.863	6.407	0.439	2.634
	Laser2	HG	89.201	2.337	17.603	73.705	214.842	6.071	0.065	2.679
		LG	87.418	0.924	18.351	34.132	204.914	5.176	0.410	2.578
245	Laser1	HG	82.292	1.988	19.121	91.821	216.201	6.827	0.113	2.680
		LG	86.137	0.868	20.354	53.562	206.207	6.403	0.378	2.613
	Laser2	HG	89.212	2.499	17.250	72.115	214.742	6.014	0.053	2.672
		LG	87.423	0.903	18.384	38.746	204.818	5.173	0.406	2.575
246	Laser1	HG	82.193	2.021	19.412	99.766	216.500	6.744	0.138	2.690
		LG	85.883	0.885	20.696	58.725	206.525	6.252	0.452	2.632
	Laser2	HG	88.588	2.447	17.740	80.218	215.222	6.104	0.074	2.681
		LG	86.821	0.913	18.735	37.230	205.306	5.242	0.439	2.591
282	Laser1	HG	82.701	1.984	19.278	95.315	216.388	6.834	0.132	2.685
		LG	86.074	0.865	20.544	55.781	206.394	6.40	0.426	2.619
321	Laser1	HG	82.43	2.071	19.015	93.349	216.230	6.789	0.116	2.680
		LG	86.144	0.834	20.629	54.728	206.231	6.369	0.429	2.616
	Laser2	HG	88.838	2.532	17.506	78.394	215.138	6.064	0.065	2.678
		LG	87.228	0.953	18.465	36.427	205.216	5.199	0.416	2.592
337	Laser1	HG	82.22	1.947	19.429	96.795	216.400	6.803	0.139	2.688
		LG	86.016	0.823	20.853	56.763	206.412	6.344	0.452	2.624
	Laser2	HG	88.564	2.496	17.691	80.859	215.287	6.104	0.069	2.685
		LG	87.078	0.944	18.614	37.446	205.380	5.235	0.426	2.584
344	Laser1	HG	82.52	2.040	18.966	90.768	216.073	6.775	0.139	2.679
		LG	86.089	0.895	20.242	53.214	206.036	6.311	0.452	2.617
	Laser2	HG	88.955	2.534	17.306	74.469	214.882	6.045	0.069	2.676
		LG	87.273	0.942	18.315	34.623	204.960	5.176	0.426	2.583
358	Laser1	HG	82.364	2.092	18.901	92.026	216.129	6.798	0.110	2.675
		LG	86.161	0.861	20.417	53.823	206.118	6.369	0.416	2.618
	Laser2	HG	88.942	2.528	17.378	75.631	215.007	6.052	0.060	2.674
		LG	87.265	0.909	18.539	35.226	205.089	5.209	0.421	2.586

HG*: High gain mode; LG*: Low gain mode

Table 3. The calculation results of the transmitted waveform during different time and orbits



(a) Variation of the mean value of BN

(b) Variation of the standard deviation of BN

(c) Variation of the mean value of SNR

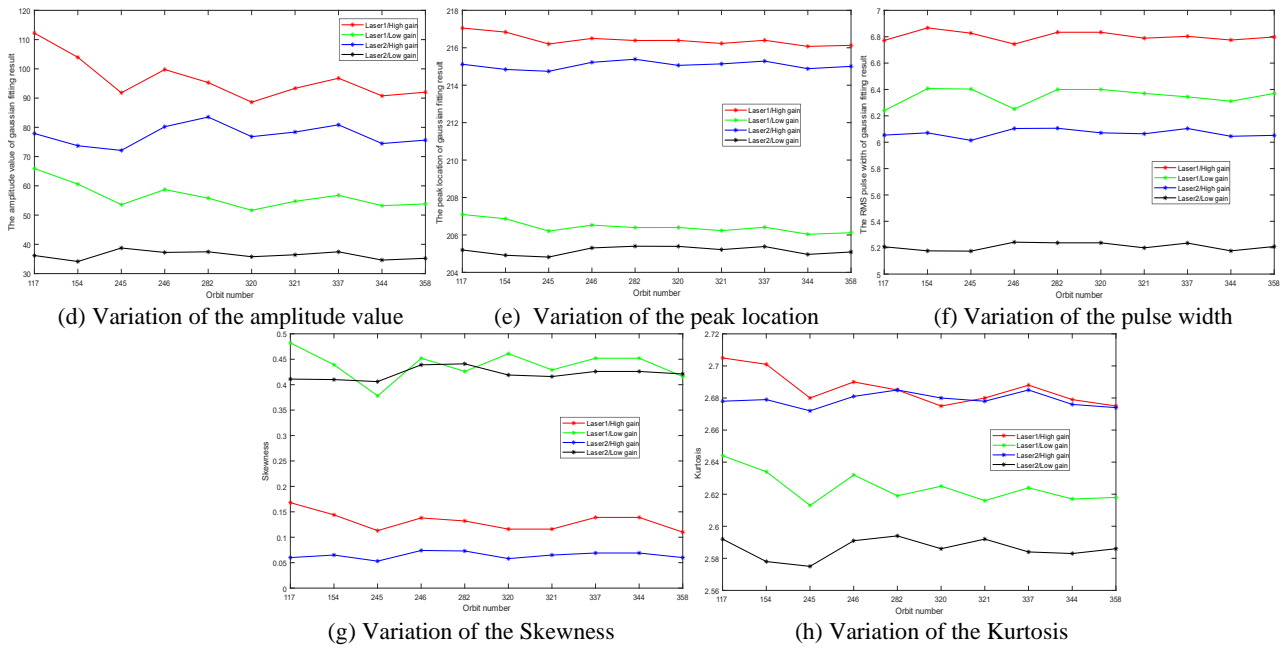


Figure 2. Analysis results of quality parameters of transmitted waveform

4.2 Experiment results of received waveform

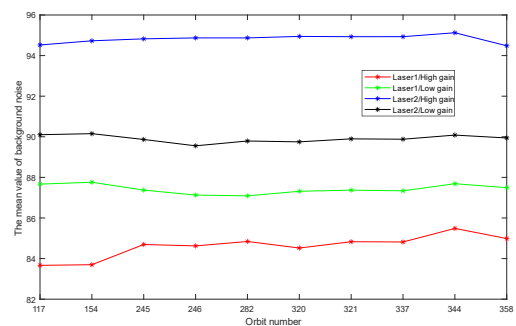
The calculated result of the background noise parameters and SNR of the received waveform of the 10 orbits data has been showed as Table 4 and Figure 3.

Orbit Number	Laser	Gain mode	Background noise(mV)		Mean of SNR
			Mean	Std	
117	Laser1	HG*	83.667	1.935	24.609
		LG*	87.669	0.886	25.137
	Laser2	HG	94.523	2.628	22.269
		LG	90.102	0.956	22.868
154	Laser1	HG	83.698	1.886	24.437
		LG	87.760	0.829	25.096
	Laser2	HG	94.728	2.337	21.820
		LG	90.154	0.924	22.133
245	Laser1	HG	84.694	1.988	24.826
		LG	87.372	0.868	24.307
	Laser2	HG	94.825	2.499	22.204
		LG	89.867	0.903	21.559
246	Laser1	HG	84.625	2.021	24.126
		LG	87.129	0.885	24.789
	Laser2	HG	94.871	2.447	22.890
		LG	89.555	0.913	22.533
282	Laser1	HG	84.842	1.984	25.929
		LG	87.091	0.865	27.993
	Laser2	HG	94.871	2.383	25.309
		LG	89.792	0.911	25.846
320	Laser1	HG	84.520	2.055	25.617
		LG	87.316	0.981	25.984
	Laser2	HG	94.945	2.535	24.174
		LG	89.750	0.934	23.603
321	Laser1	HG	84.830	2.071	25.356
		LG	87.370	0.834	25.574
	Laser2	HG	94.930	2.532	23.118
		LG	89.898	0.953	22.260
337	Laser1	HG	84.818	1.947	25.020
		LG	87.339	0.823	24.724
	Laser2	HG	94.932	2.496	22.588
		LG	89.880	0.944	21.560
344	Laser1	HG	85.486	2.040	22.659
		LG	87.688	0.895	21.309

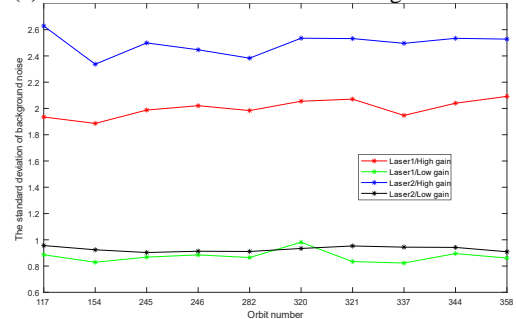
Laser2	HG	95.125	2.534	19.998	
	LG	90.083	0.942	18.865	
Laser1	HG	84.986	2.092	23.799	
	LG	87.493	0.861	23.680	
358	Laser2	HG	94.483	2.528	21.621
		LG	89.942	0.909	21.133

HG^{*}: High gain mode; LG^{*}: Low gain mode

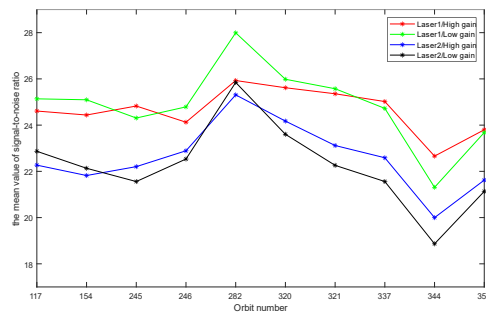
Table 4. The statistical result of the received waveform



(a) Variation of the mean value of background noise



(b) Variation of the standard deviation of background noise



(c) Variation of the mean value of signal-to-noise ratio
Figure 3. Analysis results of quality parameters of received waveform

4.3 Statistics of the APD temperature

The real-time temperature of APD has been recorded in GF-7 laser altimeter, which can be used to evaluate the state of the laser. In this paper, the relationship between the APD temperature and SNR of the transmitted waveform is illustrated in Figure 4. It is clear that SNR of the transmitted waveform can drop or decrease as the increase of APD temperature.

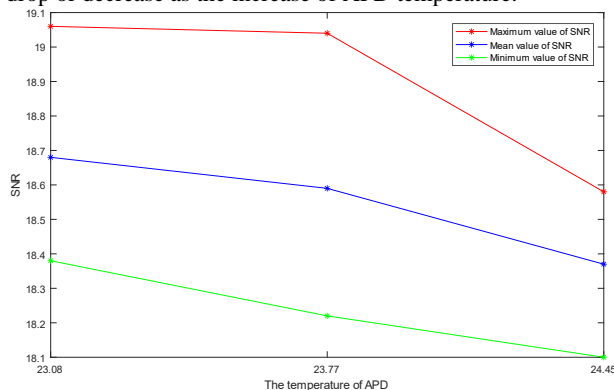


Figure 4. SNR of the transmitted waveform data at different APD temperatures (taking laser-2 as an example)

5. SUMMARY AND PROSPECT

Through the analysis of the waveform data obtaining during the period of on-orbit test of GF-7, the conclusions can be drawn as follows,

- 1) For the transmitting waveform, the mean value of the background noise of waveform varies in a similar trend between the high and low gain of the same laser, and the background noise of the waveform in high-gain mode is higher than that of low-gain mode. There is a correlation between the standard deviation of the background noise of waveforms in same gain mode of different lasers.
- 2) After the gaussian fitting of the transmitted waveform, the amplitude parameter has a most obvious variation among the gaussian parameters, and the maximum difference between the amplitude of the waveform in same gain of same laser that in different orbit is approaching 20, which mean the energy is variable. The peak location of the waveform has a maximum stability, whether in the same orbit or not, and the variation is less than 0.25ns. The RMS pulse width of the waveform of the transmitted waveform is very stable in the same orbit, and varies slightly from one orbit to the other, only within 0.1ns.
- 3) The calculation results of kurtosis value of the transmitted waveform show that the peak of the GF-7 transmitted waveform is slightly flat compared with the standard gaussian waveform, whereas, is not far from the standard gaussian. The kurtosis value is greater than 2.5. The mean of the skewness of waveform of

every orbit is with 0~0.5, indicating that the transmitted waveform is slightly skew to the right, and there is a good symmetry of the waveform yet.

- 4) In terms of the signal-to-noise ratio, the calculation method of SNR in this paper relies on the maximum intensity value of the waveform sampling point, while the maximum intensity value of the transmitted waveform is relatively small due to the different gain, thus, the SNR of the transmitted waveform is smaller than that of the received waveform. The SNR of the received waveform is affected by the terrain features, such as clouds, trees and vegetation, and the maximum of SNR of received waveform is 27.993, and the minimum value is 18.865, which shows a good reliability of the received waveform even it has the lower SNR.
- 5) The temperature of the APD will affect the quality of the transmitted waveform data, which can be used as an indicator to evaluate the state of the laser.

Quality evaluation and change monitoring

As the first laser altimeter for operational application of China, the GF-7 satellite laser altimeter with full waveform data can represent a new-era, although it has some distance with the GLAS and GEDI. After the on-orbit test, the GF-7 laser altimeter can be used to obtain the elevation control points for the stereo camera to realize the 1:10000 scale mapping, which will be exciting for the global mapping.

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