

Radiometric Cross-calibration of KOMPSAT-3A with Landsat-8

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ABSTRACT:

In this study, Cross calibration was conducted at the Libya 4 PICS site on 2015 using Landsat-8 and KOMPSAT-3A. Ideally a cross calibration should be calculated for each individual scene pair because on any given date the TOA spectral profile is influenced by sun and satellite view geometry and the atmospheric conditions. However, using the near-simultaneous images minimizes this effect because the sensors are viewing the same atmosphere. For the cross calibration, the calibration coefficient was calculated by comparing the at sensor spectral radiance for the same location calculated using the Landsat-8 calibration parameters in metadata and the DN of KOMPSAT-3A for the regions of interest (ROI). Cross calibration can be conducted because the satellite sensors used for overpass have a similar bandwidth. However, not all satellites have the same color filter transmittance and sensor reactivity, even though the purpose is to observe the visible bands. Therefore, the differences in the RSR should be corrected. For the cross-calibration, a calibration coefficient was calculated using the TOA radiance and KOMPSAT-3 DN of the Landsat-8 OLI overpassed at the Libya 4 Site. As a result, the accuracy of the calibration coefficient at the site was assumed to be $\pm 1.0\%$. In terms of the results, the radiometric calibration coefficients suggested here are thought to be useful for maintaining the optical quality of the KOMPSAT-3A.

1. INTRODUCTION

KOrea Multi-Purpose SATellite (KOMPSAT)-2 and -3 were commissioned to satisfy the demand for high-resolution satellite images in Korea. On March 26, 2015, KOMPSAT-3A was launched successfully, and Calibration and Validation (Cal/Val) began recently to improve image quality during the Launch and Early Operation Phase (LEOP). Cal/Val is divided into three parts: geometric, radiometric, and spatial. In this paper, we report the results of the radiometric Cal/Val of the spectral application images.

Since launching the KOMPSAT series, their imagery has been radiometric characterized by the Korea Aerospace Research Institute (KARI) and the Pukyong National University Remote Sensing Group (PKNU RSG) (Kim et al. 2015; Lee et al. 2012). In this paper, post-launch in-flight calibration was conducted to promote the stable operation of KOMPSAT-3A during the LEOP. The main contributions of this paper are Cross calibration between the two sensors using simultaneous image pairs, acquired by these two sensors over the Libya 4 PICS site. Cross calibration was executed based on the Landsat-8 OLI. Also, the differences in spectral responses between KOMPSAT-3 and Landsat-8, we applied the SBAF using the EO-1 Hyperion.

2. MEHTOD

In this study, Cross calibration was conducted at the Libya 4 PICS site on 2015, using Landsat-8 and KOMPSAT-3 Three times (Figure 2 and Table 1). Ideally a cross calibration should be calculated for each individual scene pair because on any given date the TOA spectral profile is influenced by sun and satellite view geometry and the atmospheric conditions. However, in this study we using the 2~3 days different images because KOMPSAT-3A is just lunched and LEOP Phase.

Site	Sensor	Date	Time (UTC)	Sun		Satellite	
				Zenith (°)	Azimuth (°)	Zenith (°)	Azimuth (°)
Libya 4	K3A	15/04/25	11:41	24.1	233.5	10.2	256.9
	L8	15/04/22	08:54	26.6	124.1	nadir	
	K3A	15/07/09	11:35	16.2	250.5	6.0	76.3
	L8	15/07/11	08:54	22.6	101.9	nadir	
	K3A	15/07/27	11:37	17.9	240.5	3.3	76.6
	L8	15/07/27	08:54	23.9	108.3	nadir	

Table 1. Metadata of images in this study

For the cross calibration, the calibration coefficient was calculated by comparing the at sensor spectral radiance for the same location calculated using the Landsat-8 calibration parameters in metadata and the DN of KOMPSAT-3A for the regions of interest (ROI).

Each region of interest (ROI) for Cross calibration (Libya 4 Site) was selected to be homogenous, with a size of about $9 \text{ km} \times 9 \text{ km}$. To maintain equivalence in the pixel size between KOMPSAT-3 and Landsat-8 images, the KOMPSAT-3A images were resampled to a 30-m pixel size from their original 2.4-m resolution for multispectral bands using bilinear resampling.

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2.1 Sensor Overview

Table 1. lists the key specifications of the KOMPSAT-3A AEISS-A, Landsat-8 OLI, and the EO-1 Hyperion sensor. Only the four bands that the three sensors have in common, as listed in the table. for clarity the bands considered in this paper are named Blue, Green, Red, and Near Infrared (NIR).

Parameter	EO-1 Hyperion	KOMPSAT-3A	Landsat-8
Spectral Bands	B B11 ~18 0.52	B1 0.45~ 0.52	B2 0.45~ 0.52
	G B16 ~25 0.59	B2 0.52~ 0.60	B3 0.55~ 0.60
	R B28 ~34 0.69	B3 0.63~ 0.69	B4 0.63~ 0.68
	N B39 ~56 0.90	B4 0.76~ 0.90	B5 0.84~ 0.88
GSD at Nadir	30 m	2.8 m	30 m
Quantization	10 bit	14 bit	12 bit
Swath width	7.5 km	15 km	185 km
Altitude	705 km	685 km	705 km
Scanner	Push broom	Push broom	Pushbroom

Table 2. Metadata of images in this study

Landsat-8 is the latest platform in the 40 year Landsat series of satellites. The visible and near infrared (VNIR) bands retain many of the characteristics of previous Landsat sensors, such as a 185km swath width, a 30m ground instantaneous field of view (GIFOV) for the multispectral bands, and a 15m GIFOV for the panchromatic band. One prominent change was the transition from a whiskbroom configuration (ETM+) to a pushbroom configuration, which was successfully demonstrated by the EO-1 Advanced Land Imager (ALI) (Irons et al. 2012).

The NASA EO-1 satellite was launched on November 21, 2001 as part of a one year technology validation/demonstration mission. The Hyperion pushbroom hyperspectral sensor covers the 0.4–2.5µm spectral range with 242 spectral bands at a 10nm spectral resolution and a 30m spatial resolution, acquiring spectra over a 7.7km swath width. The spectral characterization of Hyperion is well documented (Ungar et al. 2009), and the Hyperion RSR is characterized using a Gaussian profile. Overall, the EO-1 Hyperion sensor is radiometrically stable to within 5%, based on the lunar observations in comparison with the United States Geological Survey (USGS) Robotic Lunar Observatory (ROLO) (Folkman et al. 2001).

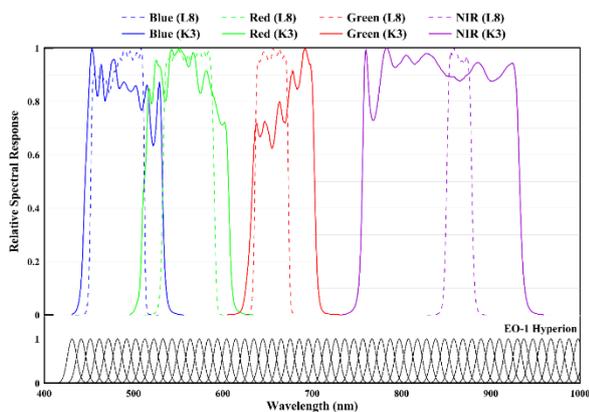


Figure 1. RSR of the sensor used in this study

Figure 2 shows the band average RSR of KOMPSAT-3A together with the Landsat-8 OLI and EO-1 Hyperion. In general, KOMPSAT-3 bands are narrower than the Landsat-8 OLI bands because the Landsat-8 OLI band edges have been refined to avoid atmospheric absorption features. The Landsat-8 OLI NIR band is substantially narrower, and so avoids the water vapor absorption feature at approximately 825, 940 nm. These differences mean that even when both sensors are observing the same general region of the electromagnetic spectrum at the same time, they may report different radiance values depending on the spectral signature of the target.

2.1 Libya-4 Site

PICS-based calibration is only good as long as the sites being used continue to exhibit a reasonable spatial, temporal and spectral stability. Optimal test sites for radiometric calibrations are spectrally and spatially homogenous, vary little over time, are near-Lambertian, and have higher surface reflectance. Therefore, PICS are typically used for EOS radiometric cross-calibration (Dingirard and Slater 1999). In particular, deserts provide optically stable sites suitable for absolute sensor calibration, for example in Libya, Algeria, Niger, and Mauritania. This study cross-calibration TOA reflectance with the Landsat-8 OLI sensor, a multispectral sensor with similar bandwidth, to evaluate the vicarious calibration coefficient of KOMPSAT-3A AEISS-A sensor radiometric calibration.

Libya 4 is a high-reflectance site located in the Libyan Desert in Africa at coordinates +28.55° N and +23.39° E and at an elevation of 118 m above sea level. It is referenced in the worldwide reference system 2 with path 181 and row 40. The site is named Libya 4 as opposed to Libya because of the presence of other pseudo invariant sites in Libya. Figure 2 shows the Libya-4 PICS true color image of KOMPSAT-3A and Landsat-8 used for cross calibration.

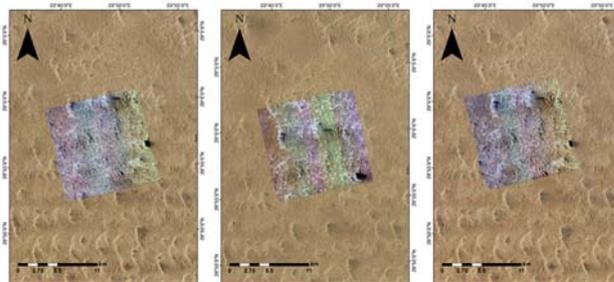


Figure 2. K3A and L8 Libya-4 Site Image for Cross Calibration

2.2 TOA Radiance and Reflectance

The Landsat-8 OLI uses two independent National Institute for Standards and Technology (NIST), USA traceable radiance and reflectance calibration methods. We used the Landsat-8 metadata file, including the coefficients necessary to convert to radiance, and the reflectance from the quantized and calibrated DNs of the product (Helder et al. 2013).

This can achieve rapid conversion by calculating the radiometric rescaling factor offered in the MTL file using the following formula:

$$Radiance_{Landsat8} = DN * MULT + ADD \quad (1)$$

Where, MULT = Band-specific multiplicative rescaling factor, ADD = Band-specific additive rescaling factor, DN = Quantized and calibrated standard product pixel values

The at-sensor or TOA reflectance is used to reduce illumination differences among images and to normalize for solar irradiance. Because atmospheric effects and surface topographic effects are not corrected, this includes the Earth's surface reflectance and atmospheric reflectance (Cao et al. 2008). When comparing the radiometric quality obtained from the other sensors, the cosine effect of the different solar zenith angles due to the differences in time in obtaining materials could be removed if TOA reflectance was used instead of the TOA radiance. The equation to calculate the TOA Reflectance (ρ) of KOMPSAT-3 and Landsat-8 is as follows:

$$\rho_{\lambda} = \frac{\pi \cdot L_{\lambda} \cdot d^2}{ESUN_{\lambda} \cdot \cos\theta_s} \quad (2)$$

Where, ρ_{λ} is planetary reflectance, L_{λ} is spectral radiance at the sensor aperture (either KOMPSAT-3 or Landsat-8 ($W/m^2 \mu m sr$)), $ESUN_{\lambda}$ is the band dependent mean solar exoatmospheric irradiance ($W/m^2 \mu m$), θ_s is the solar zenith angle (radians), and d is the earth sun distance (astronomical units).

2.1 SBAF (KOMPSAT-3)

EO-1 Hyperion sensor images were used to derive the spectral band adjustment factor (SBAF) to compensate for the differences in the RSR between the sensors. Although SBAF is discussed briefly in this section, readers are directed elsewhere for the detailed mathematical expressions for SBAF (Chander et al. 2013).

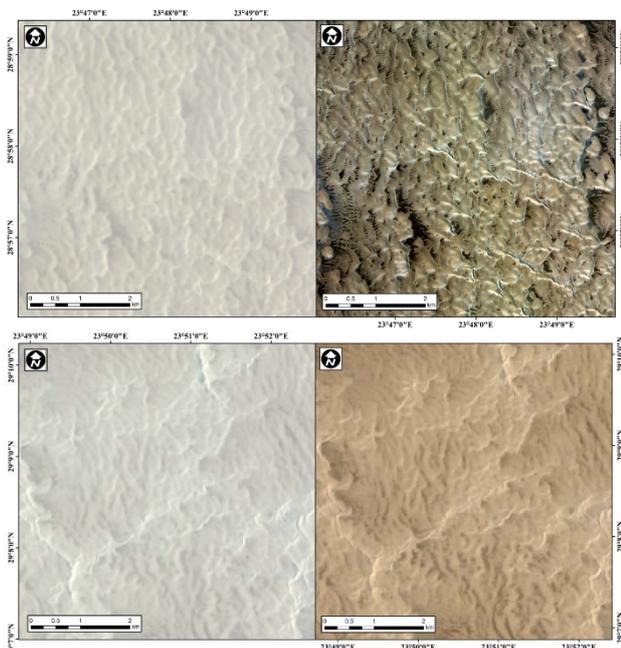


Figure 3. Common area of KOMPSAT-3, Landsat-8 and Hyperion images in Libya-4 PICS site for Applying SBAF.

The suitability of the Hyperion sensor for the assessment of spectral band differences has also been addressed elsewhere in the literature (Chander et al. 2009). The SBAF was calculated using KOMPSAT-3 and EO-1 on July 6 2014, and using Landsat-8 and EO-1 on July 8 2014 (Figure 3). SBAF can be calculated using the following formula by utilizing the integral values of the RSR [11]:

$$\overline{\rho_{\lambda}(sensor)} = \frac{\int \rho_{\lambda} * RSR_{\lambda} d\lambda}{\int RSR_{\lambda} d\lambda}$$

$$SBAF = \frac{\overline{\rho_{\lambda}(A)}}{\overline{\rho_{\lambda}(B)}} \quad \overline{\rho_{\lambda}(A^*)} = \frac{\overline{\rho_{\lambda}(A)}}{SBAF} \quad (3)$$

Here, RSR_{λ} is the RSR of the sensor, ρ_{λ} is the hyperspectral TOA reflectance profile, $\overline{\rho_{\lambda}(A)}$ is the simulated TOA reflectance for sensor A, $\overline{\rho_{\lambda}(B)}$ is the simulated TOA reflectance for sensor B, $\overline{\rho_{\lambda}(A^*)}$ is the compensated TOA reflectance for sensor A when using the SBAF to match sensor B TOA reflectance.

3. RESULT

For the cross calibration, the calibration coefficient was calculated by comparing the at sensor spectral radiance for the same location calculated using the Landsat-8 calibration parameters reported in various studies and the TOA Reflectance of KOMPSAT-3 for the target area.

Cross calibration results were tested for a slope value of 1.0 (corresponding to an exact agreement in radiances between the two sensors) at a 0.001 significance level after fitting a regression line to data from the common area of interest. For KOMPSAT-3, the regression lines were determined for data within each band, as was an overall regression combining all of the data points from all bands. For the quantitative or cross calibration of the different sensors, in which spectral resolution and characteristics differed, the pixels were converted into radiance at the sensor.

Date	Band	TOA Reflectance		Cross Coef.
		K3A	L8	
04/25 (K3) 04/22 (L8)	B	0.245	0.259	0.9459
	G	0.326	0.325	1.0031
	R	0.468	0.452	1.0354
	N	0.581	0.513	1.1326
07/09(K3) 07/11 (L8)	B	0.251	0.240	1.0458
	G	0.338	0.318	1.0629
	R	0.473	0.428	1.1051
07/27(K3, L8)	N	0.601	0.500	1.2020
	B	0.251	0.248	1.0121
	G	0.335	0.323	1.0372
	R	0.466	0.432	1.0787
	N	0.596	0.488	1.2213

Table 3. Cross Calibration Result

The TOA reflectance of Landsat-8 and KOMPSAT-3A were compared for the calibration coefficient and the percentage difference was calculated using the Libya-4 PICS (Table 3). From this, the blue and green bands were within 4%. In the red band, the TOA reflectance of KOMPSAT-3A and Landsat-8 differed by ~8%. However, in the NIR band, the average difference in TOA reflectance was 22% for the Libya site. This can be explained by the values in the band RSR of two sensors,

which differed significantly from the NIR band and, unlike Landsat-8, the KOMPSAT-3A sensor reacted in a spectral region of 940 nm that is highly affected by water vapor absorption. This explains the lower TOA reflectance of KOMPSAT-3A in the NIR band than that of Landsat-8 (Brooke 2012). However, after comparing the KOMPSAT-3 sensor with the same RSR with Spectral Band Adjustment Factor (SBAF)(Shin et al, 2015), the difference was calculated as < 5%, which was considered to be appropriate(Table 4).

Band	L8 a	K3A b	Diff b-a	% diff (b-a)/b	St dev.	SBAF (K3)	SBAF K3A (c)	Diff c-a	SBAF % diff
B	0.2510	0.248	-0.003	-1.19	0.004	0.979	0.243	-0.008	-3.27
G	0.3350	0.323	-0.012	-3.65	0.007	1.014	0.328	-0.007	-2.23
R	0.4660	0.432	-0.034	-7.78	0.011	1.023	0.442	-0.024	-5.16
N	0.5960	0.488	-0.108	-22.04	0.015	1.221	0.596	0.000	-0.03

Table 4. Apply SBAF for 07/27 Cross calibration result

4. CONCLUSION

In this study, with KARI and PKNU RSG, the relationship between the DN and sensor radiance was calculated using the radiometric calibration coefficient through cross calibration of the KOMPSAT-3A sensor.

In the cross-calibration, Standard deviation between Cross calibration coefficients are shown up to about 3% difference. Also it showed that over time, a coefficient value decreased. Therefore, it is determined that such sensor degradation phenomenon need to be analyzed. Before and after SBAF compensation in the blue and green bands, the differences in the red band improved from -7 to -5%, and the differences in the NIR band improved from -22 to 3%.

This likely resulted in the band RSR difference, as the difference between Landsat-8 and KOMPSAT-3A NIR band is due to the water vapor region in the NIR band. They are believed to provide a good scale for maintaining the optical quality of KOMPSAT-3A AEISS-A, which lacks data for prelaunch, onboard, and vicarious calibration. However, we need to analyze the nonlinearity generated by determining the calibration coefficient and reduce the uncertainty (atmosphere and BRDF effect).

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